Study of TGEs and Gamma-Flashes from Thunderstorms in 20-3000 keV Energy Range with SINP MSU Gamma-Ray Spectrometers

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A number of experiments with scintillator gamma-spectrometers were provided by SINP MSU for study of spectral, temporal and spatial characteristics of TGEs as well as for search of fast hard x-ray and gamma-ray flashes probably appearing at the moment of lightning. The measurements were done in Moscow region and in Armenia at Aragatz mountain.

TEPA-2015, 07.10.2015

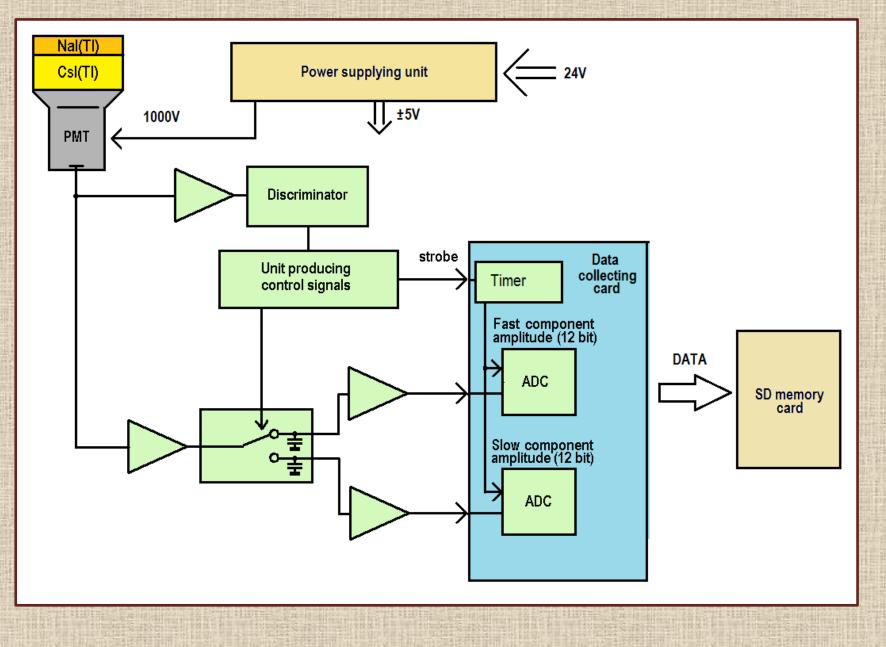
## **Scientific goal:**

- Study of spectral characteristics of TGEs in 20-3000 keV range
- Measure of the direction of TGE gamma-radiation
- Search for fast gamma-ray flashes from lightnings

### **Principles of instrument design:**

- Detectors are scintillator spectrometers with NaI(TI) or CsI(TI)
- Electronics allows to analyze pulse shape in order to use phosvich detectors and to remove imitations of gammaevents by thunderstorm electric discharges
- Recording all data in "event" mode with fine time resolution
- All data are recorded to SD card for further analysis
- Measurements with collimated detector placed on rotating platform are provided

## **Design of instrument electronics**



## **Digital electronics and data format**

- Producing time data with accuracy 1s.
  Stability of internal timer is ~1s/day
- Forming data frames each second.

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- Producing ~15mcs timer data starting at the beginning of the frame
- Interrupt on the request from analog card and digitize pulses of fast and slow components
- At the beginning of a frame digitize signal on the analog input



Board STM32F4 DISCOVERY with Cortex M4 microcontroller

7b - Frame start marker - E4 57 B4 C0 3F 66 99

- 4b Frame number
- 6b Time YY MM DD hh mm ss
- 2b Number of events in the frame
- 4b Number of counts of 15mcs timer during the frame
- 2b ADC data for external analog input
- N\*(3b+3b) Data records: ADC data + timer value
- 4b Frame end marker CC 11 00 00

### **Two gamma-ray spectrometers used in this work:**

Detector: Nal(Tl) 40x40 mm PMT: ΦЭУ-176 Range: 20 κэΒ-1 MэB Resolution 12% at 662 keV



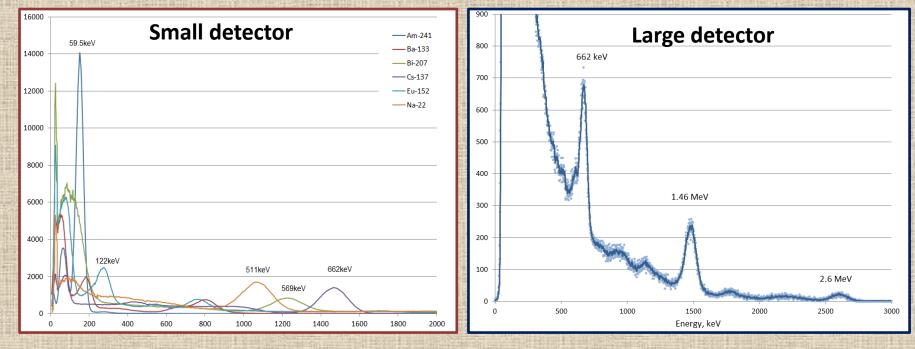
Detector: CsI(TI) 80x80 mm PMT: Hammamatsu R1307 Range: 20 кэВ-3 МэВ Resolution 7.5% at 662 keV Placed on rotating platform



## **Calibration and data processing**

# Gamma to gamma data were processed and three kinds of secondary files were produced:

- 1)Monitoring time sequences in several energy channels with 1s resolution
- 2)Detailed energy spectra for requested periods
- 3)Event data sequences (useful for short burst search)



**Autocallibration** alhorythm was used for large detector data: every 300s of the data the program determined the actual position of well visible 1.46 MeV background gamma-line of K-40, then the energy of gammas in keVs was calcullated. It allowed to minimize the effects of false variations connected with temperature drift of the detector characteristics

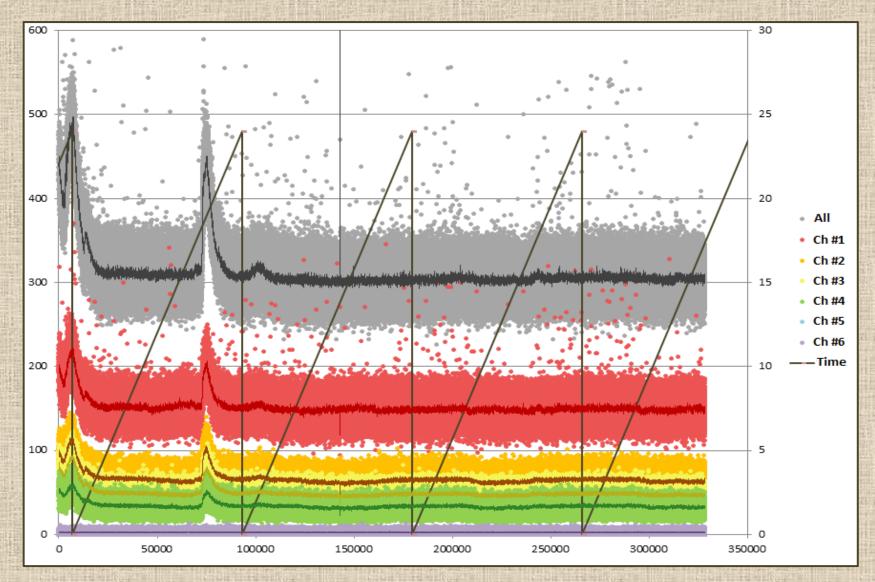
### **Observations 50 km North from Moscow**



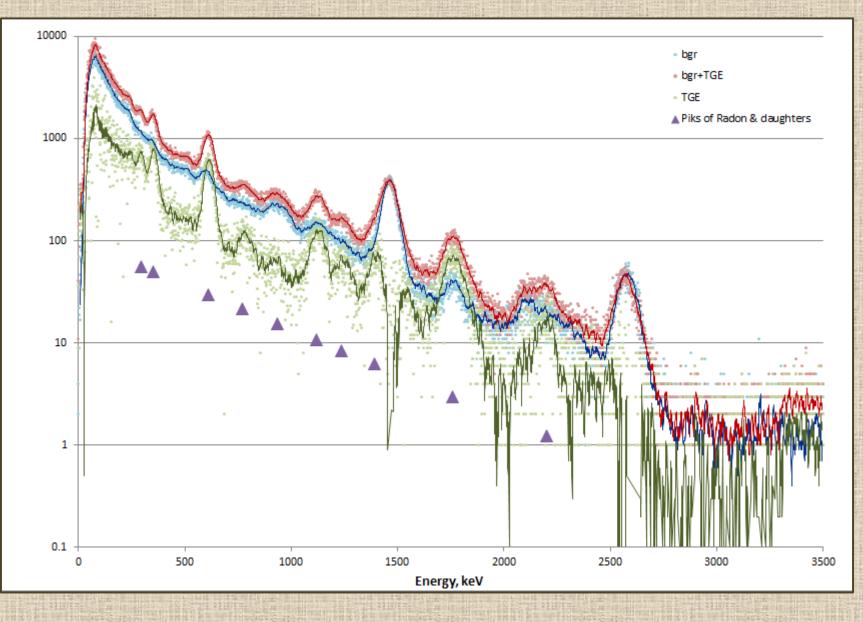




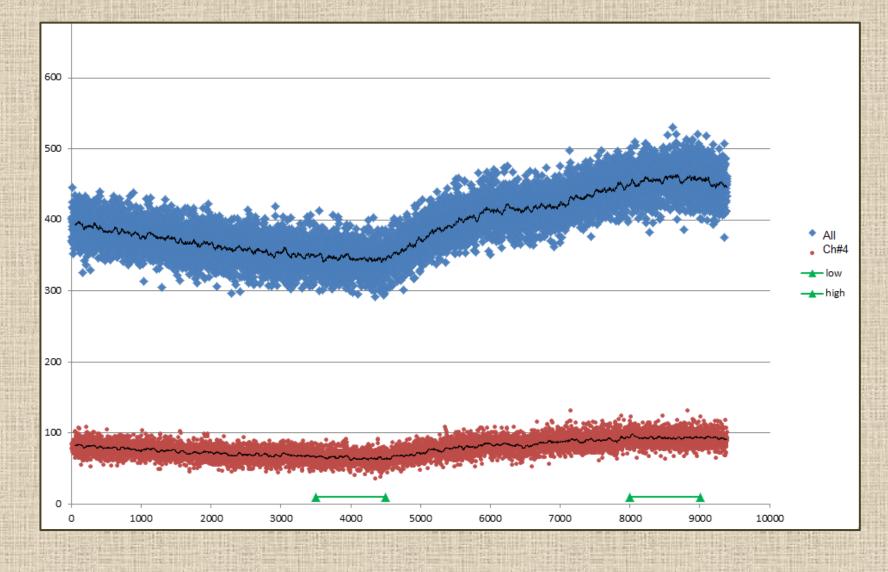
#### 27-30 of July 2015. Conditions: thunderstorms (peaks), rain, next - clear weather



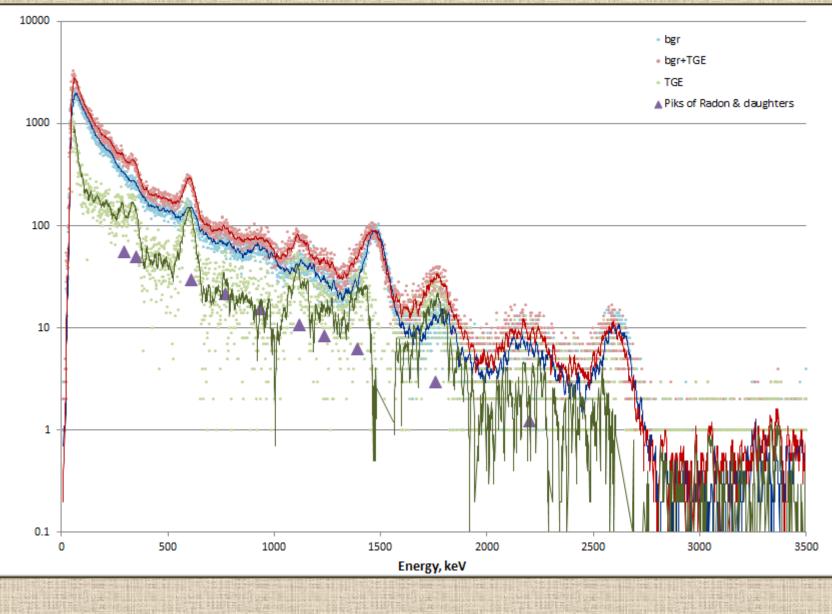
### **Energy spectrum of TGE 28.07.2015**



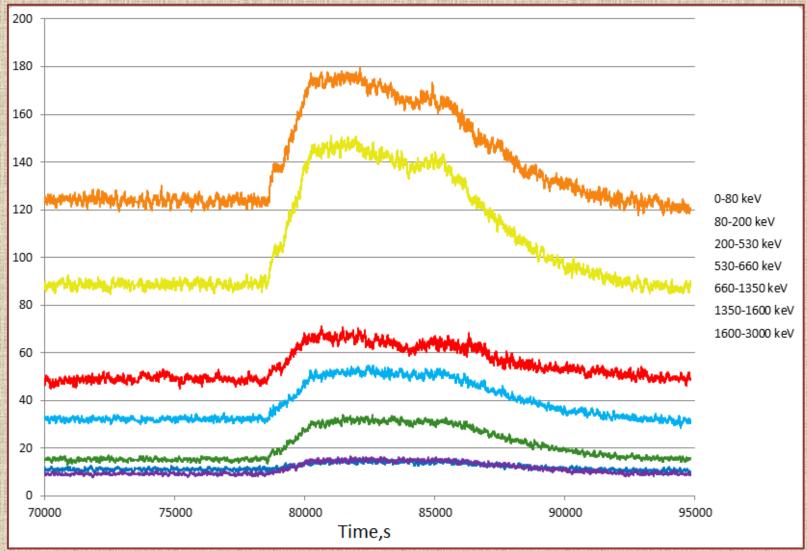
#### 05 of September 2015. Conditions: small rain without thunderstorms



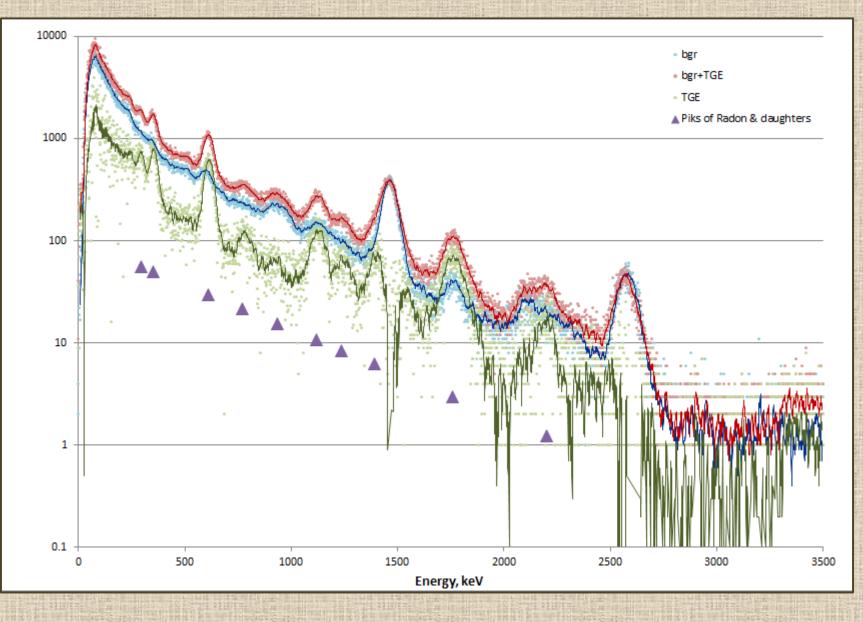
### Energy spectrum of 05.09.2015



#### 27 of September 2015. Conditions: clear weather, then thunderstorm with powerful rain, then the rain stopped and the sun appeared

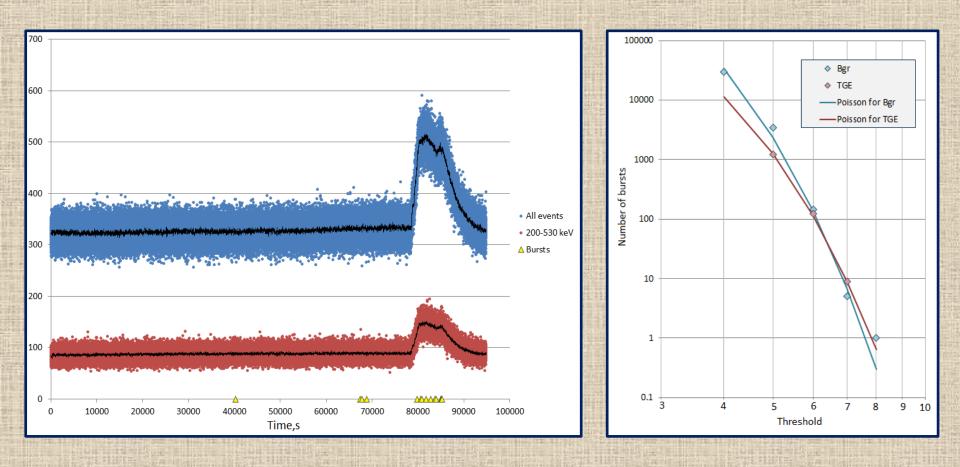


### **Energy spectrum of TGE 28.07.2015**

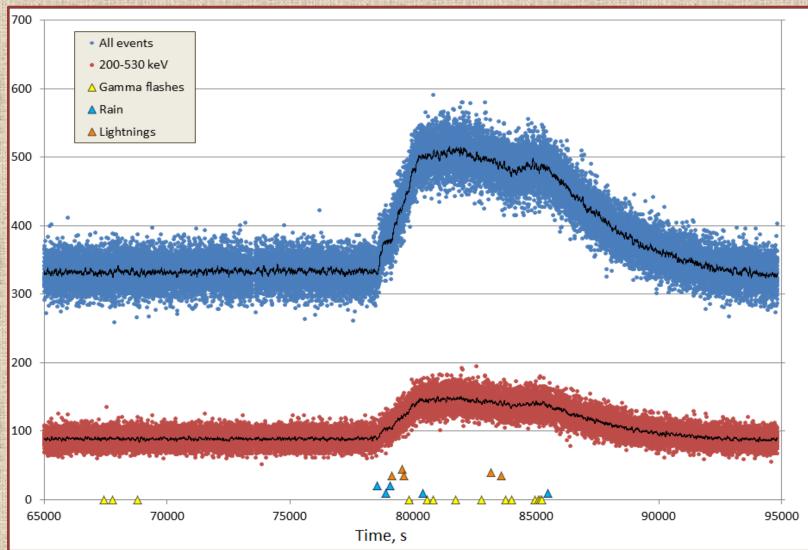


### **Search for short bursts**

The data of thunderstorm 27.09.2015 were processed. The moments of short burst candidates when >7 gammas occur in 1 ms were determined (yellow triangles) The graph of expected number of imitations vs threshold value shows that probably all candidates pointed on the left figure are random and the criterion must be some harder with threshold of 9 gammas per 1 ms. Such events were not observed.



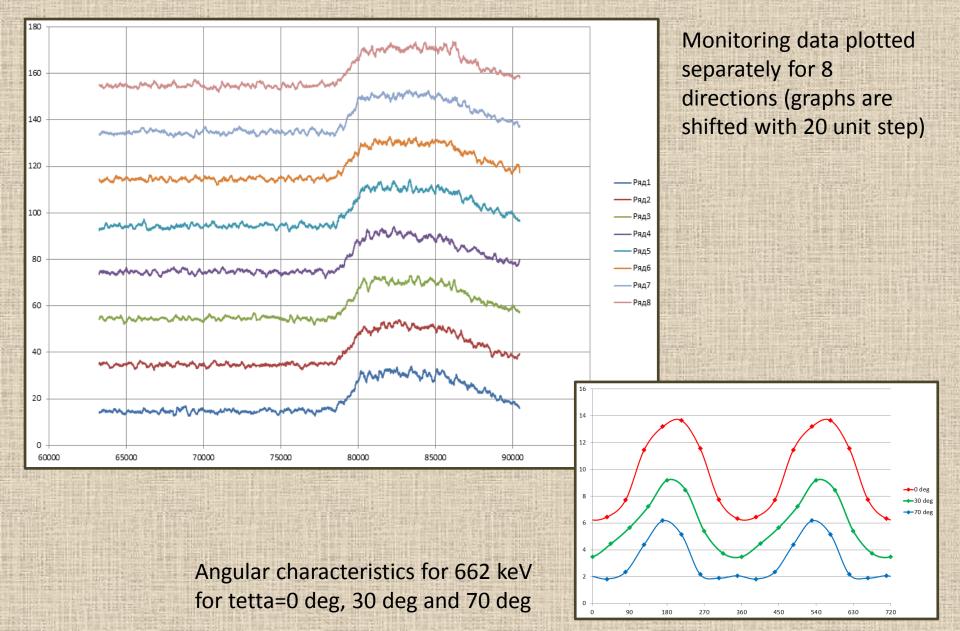
### **Comparison with weather**



1) There are no gamma-ray flashes at the moments of lightenings

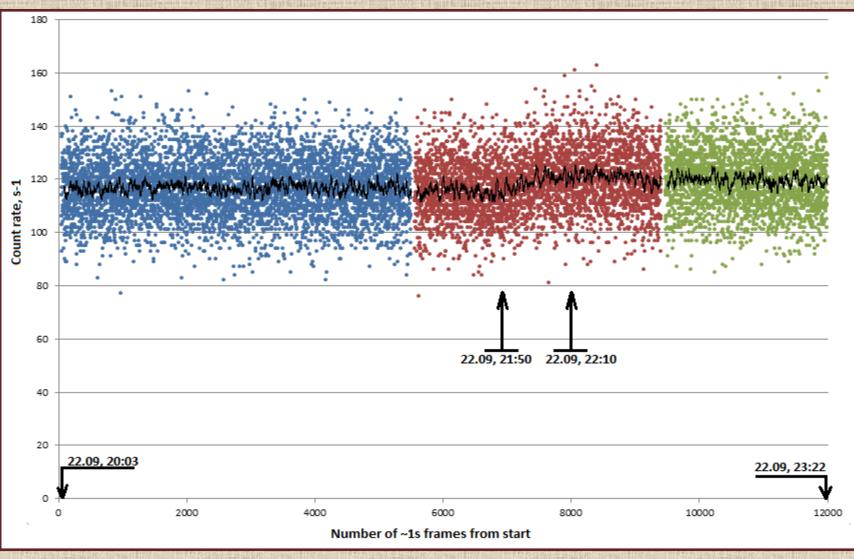
- 2) The moment of gamma-ray flux increase exactly coincide with the rain start.
- 3) When the rain stops the TGE decays exponentially with time index ~1h

### **Test for directivity**

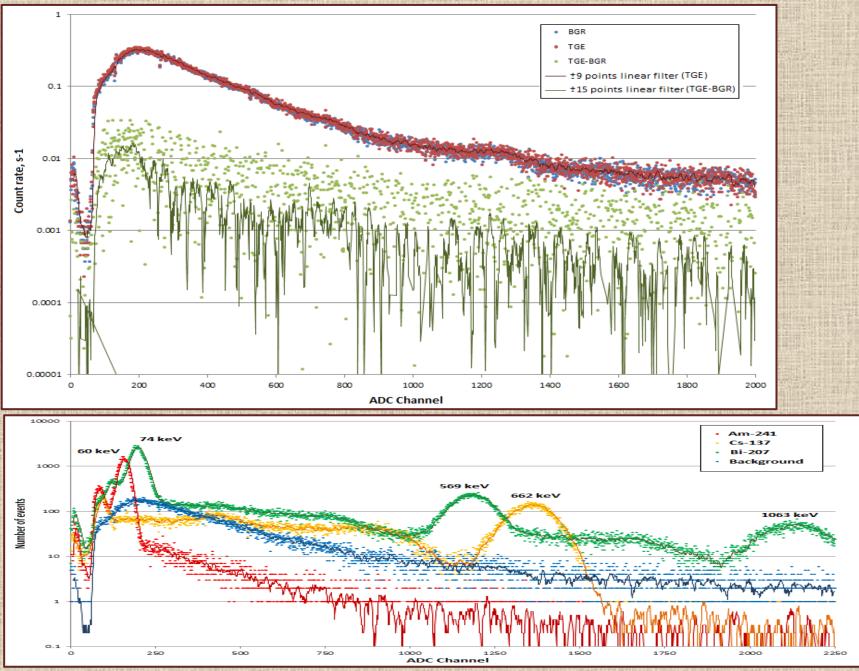


## **Measurements with Nal(TI) detector in Nor-Amberd**

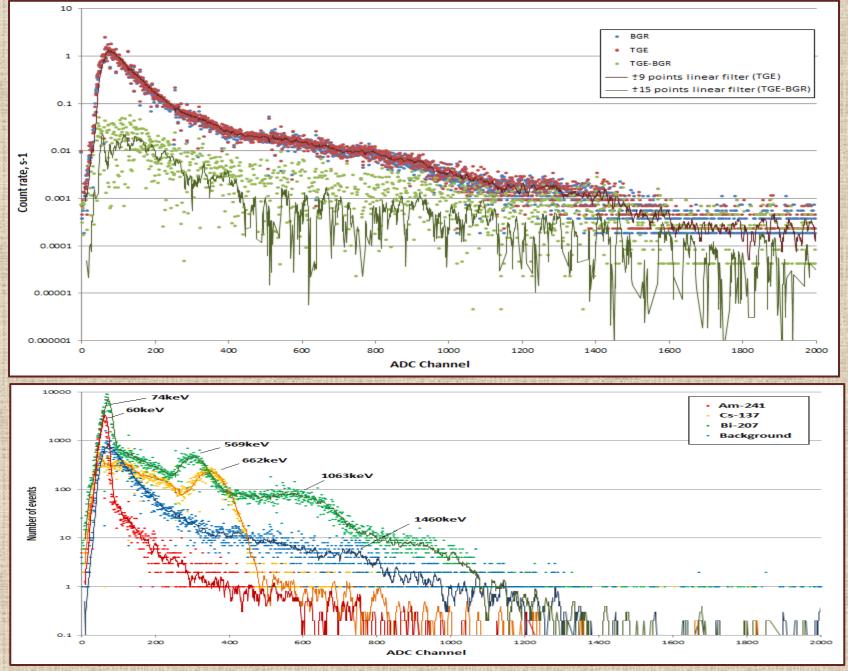
#### 22.09.2014 (during TEPA-2014)



#### Energy spectrum of TGE 22.09.2014

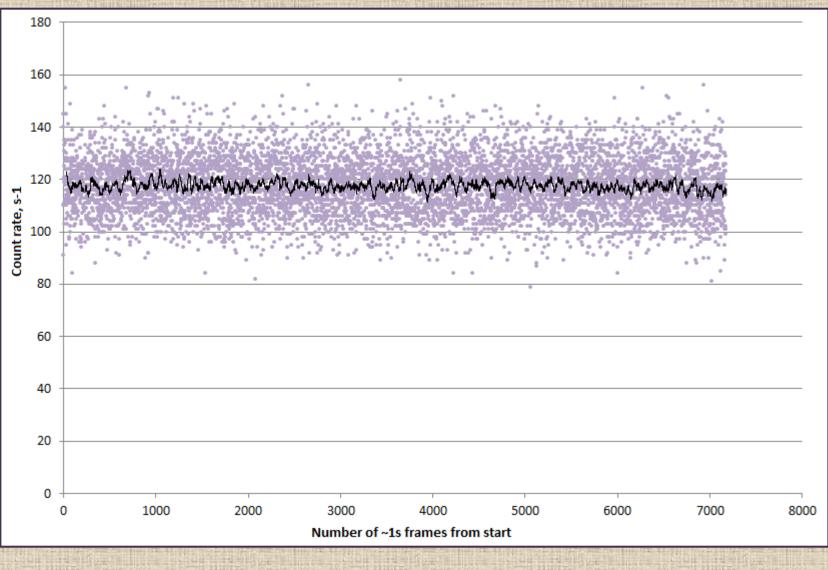


#### **Spectrum of TGE 22.09.2014 – processing for high energy range**

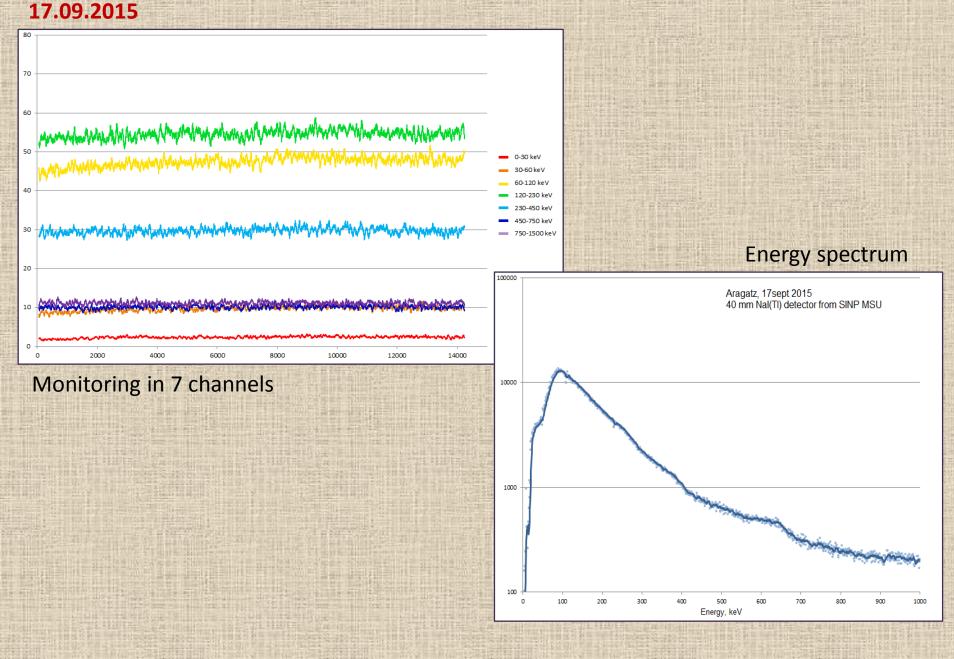


### Measurements on 3200 m

#### 24.09.2014 (during TEPA-2014)

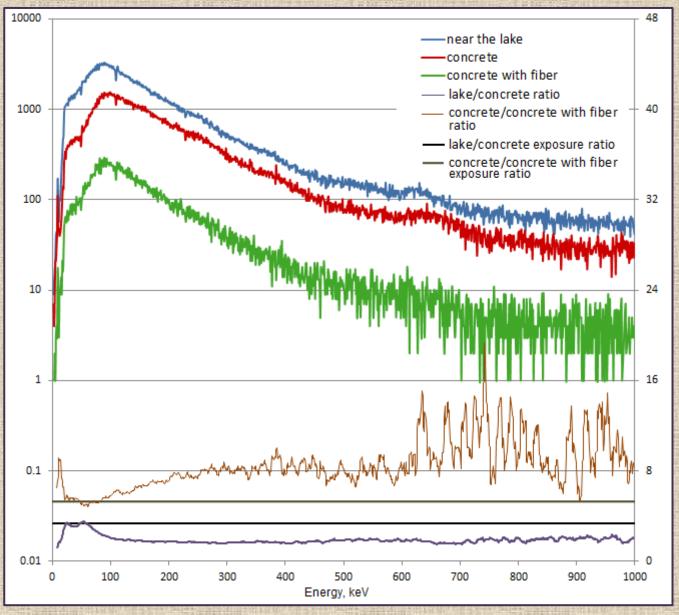


### Measurements on 3200m in 2015



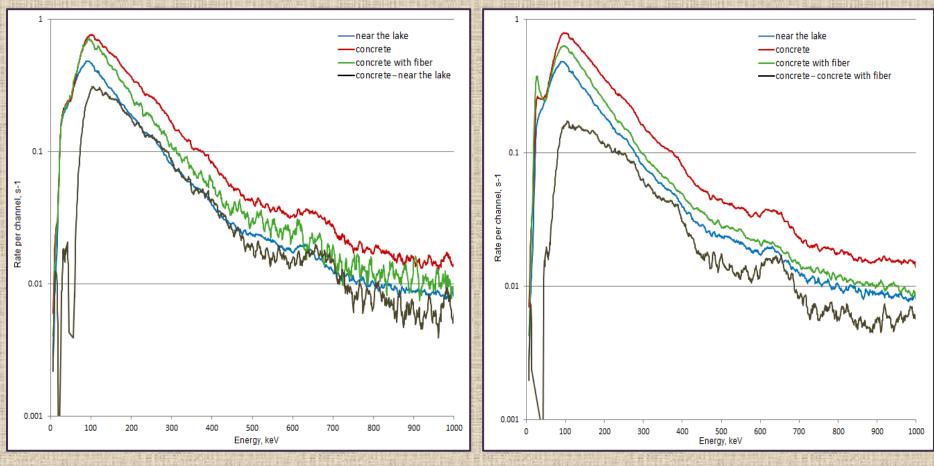
#### **Spectra at different sites of Aragatz station**

#### 20.09.2015



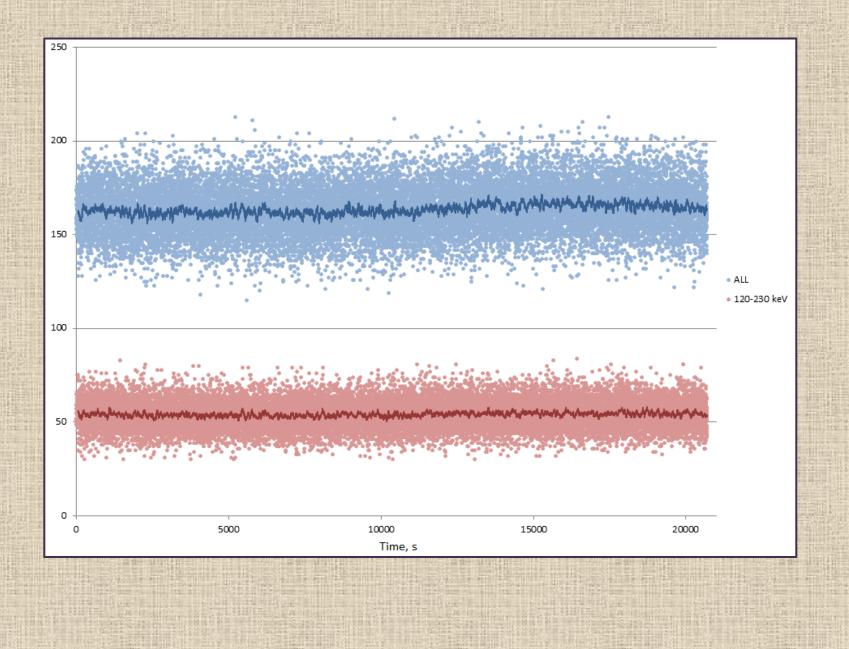
#### **Comparison of spectra at different sites of Aragatz station**





#### 22.09.2015

#### Measurements on Aragatz during TGE 04.10.2015



#### **Energy spectrum obtained during TGE 04.10.2015**

