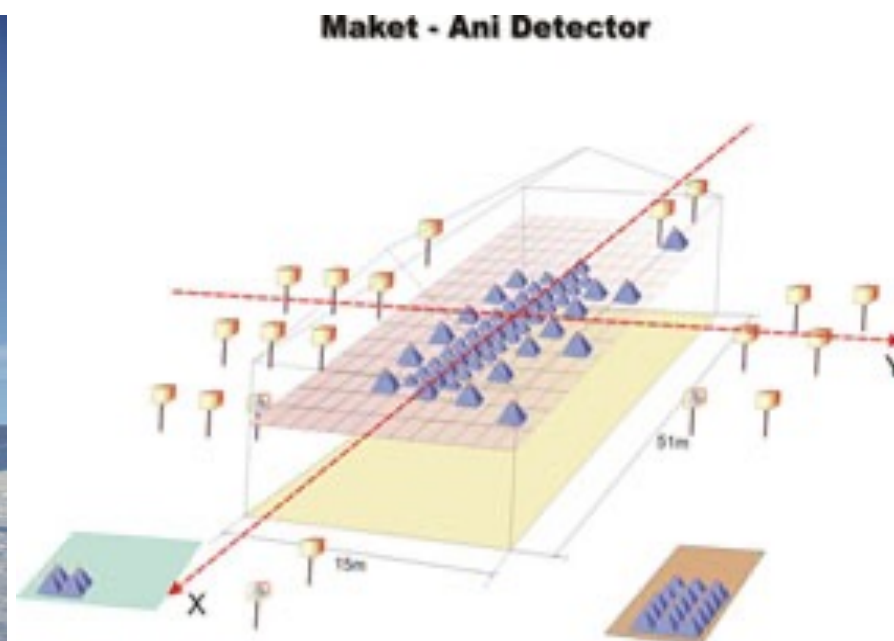


Energy spectra of light primary cosmic rays in the energy range from 10 TeV to 100 PeV

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Mt.Aragats, 3200m, 40.5°N, 44.2°E,  
92 plastic scintillators, effective area for EAS  
collection  $\sim 1000\text{m}^2$  for  $N_e > 10^5$ .



The non-parametric multivariate methodology of data analysis allows event-by-event classification of all EASs and first-time present light and heavy nuclei energy spectra separately. p+He spectrum obtained by MAKET is in good agreement with the spectra from balloon and satellite measurements, within the QGSJet-II model.

The energy spectrum of the light component (p+He) measured by the MAKET-ANI detector in comparison with the spectra published by KASCADE, EAS-TOP, HEGRA, EAS-TOP+MACRO, and TIBET. The direct balloon measurements by ATIC-2 and JACEE, HAWC, PAMELA, DAMPE, and CREAM also are presented.

The QGSJet-II and QGSJet-III models (S.Ostapchenko, ISVHECRI-2022, India) predicted 10% higher  $N_e$  and  $N_{\mu}$ . This should impact the energy reconstruction and thereby rescale down the normalization of our results, bringing them closer to most of the other measurements.

